

Apollo: Shoot the Moon

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April 14, 2015

Abstract

The objective of this mission was to have a spacecraft leave Earth's parking orbit (175 km), venture to the Moon, and return back to Earth. To do this a program with proper force laws needed to be written. It was necessary that the program also began with specific initial parameters.

1 Introduction

Here we began by modeling the Earth-Moon system using the gravitational forces of one body on the other. In order to properly model the system we had to solve for the center of mass (barycenter). The barycenter is the point that each object in our system orbits. Using Newton's Force Laws we modeled the force each object had on the other:

$$\vec{F}_{1,2} = -\vec{F}_{2,1} = \frac{GM_1M_2}{|\vec{r}_1 - \vec{r}_2|^3}(\vec{r}_1 - \vec{r}_2) \quad (1)$$

where $\vec{F}_{1,2}$ is the force of one body acting on the other, G is the Gravitational constant, M_1 is mass of object one, M_2 is the mass of object two, \vec{r}_1 and \vec{r}_2 are the respective position vectors of each body. From these standard force laws the acceleration each body caused on the other was able to be determined. This was necessary in order to implement the Runge-Kutta method of numerical analysis.

To force structure on the system, Earth was placed at the proper distance below the barycenter. The Moon's initial location was located in the plane, above the barycenter.

$$r_E = \frac{M_m}{M_m + M_E}D \quad (2)$$

$$r_m = \frac{M_E}{M_m + M_E}D \quad (3)$$

where r_E is the distance from the center of the Earth to the barycenter, r_m is the distance from the center of the Moon to the barycenter, M_m is the mass of the moon, M_E is the mass of the

Earth, and D is the Earth-Moon distance. Knowing $M_E \gg M_m$ tells us that the barycenter is located very close to the center of mass of the Earth. Once a proper Earth-Moon system was modeled we introduced a third body, the Apollo 13 spacecraft. The objective was to simulate the actual flight path of the Apollo 13 mission.

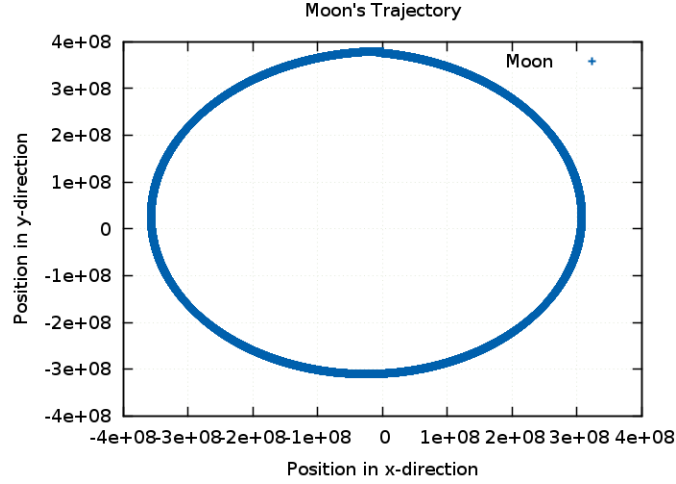


Figure 1: *The orbit of the moon, about the barycenter of the Earth-Moon system*

Figure 1 is a graphical representation of the path traced out by the Moon's orbit, about the barycenter, in one sidereal Earth-Moon orbital period of 27.322 days.

2 Computational problem

During the Apollo 13 mission an issue with the No. 2 oxygen tank forced the mission to be modified. Instead of landing on the moon the astronauts aboard were forced to make a course correction. It was this correction that led to their gravitational slingshot around the Moon. After tracing out a figure-eight flight pattern the crew was able to land back on Earth safely.

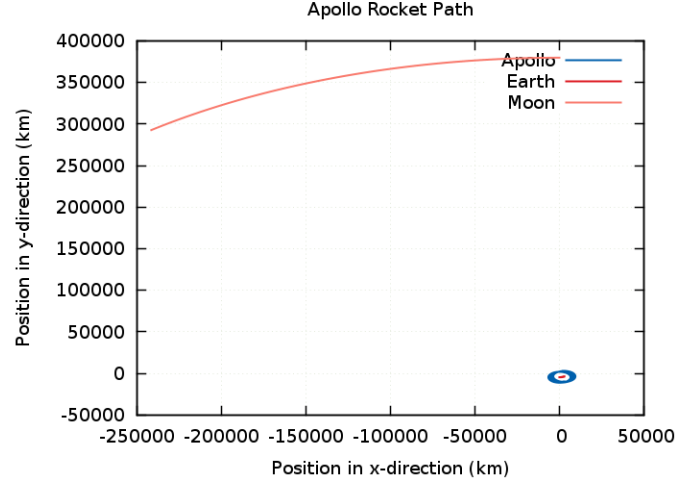


Figure 2: *The Earth-Moon system with a satellite orbiting the Earth*

Figure 2 shows the orbits of the Earth, Moon, and Apollo spacecraft in parking orbit above the Earth over a period of 3 days. In order to model the path of Apollo, as it ventured out of Earth's parking orbit, a change in velocity was necessary. After summing the velocity of the spacecraft with that of Earth's, a correction had to be made. For simplicity, the newly added velocity was considered to be instantaneous. By adding more velocity Apollo broke free of Earth's parking orbit and began its transit to the moon.

The challenge was in uncovering the correct magnitude and direction of the necessary velocity correction. When done properly the spacecraft would loop around the Moon and return the onboard astronauts safely to Earth. Computationally, this is no easy task. First an attempt at a Hohmann Transfer Orbit was made. In such a simplistic scenario this approach quickly proved to surpass our programming capabilities. In order to achieve the desired goal a trial and error method was implemented. Our first task was to hit the moon.

3 Results

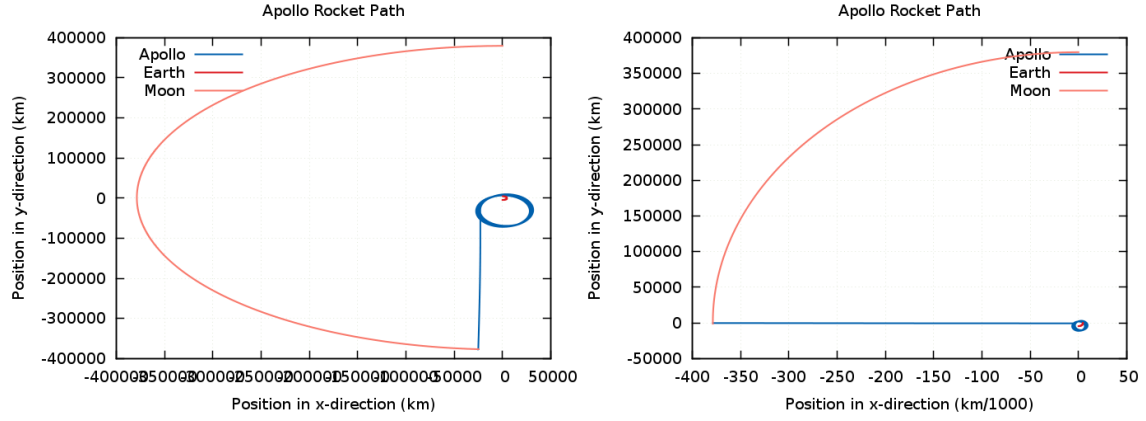


Figure 3: *Two different trajectories that cause a lunar crash*

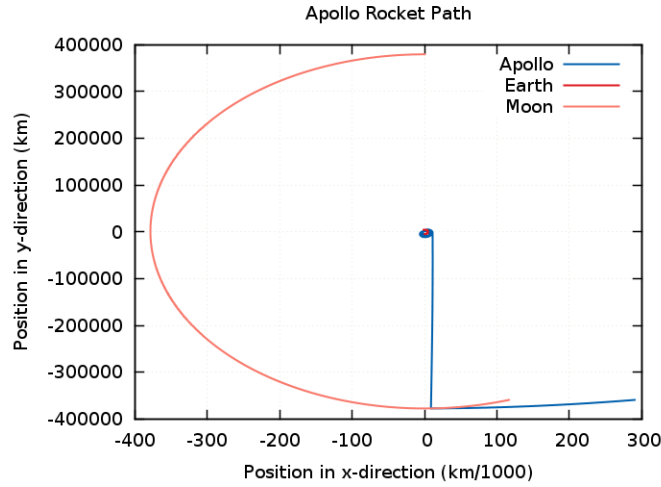


Figure 4: *The Apollo 13 spacecraft unable to orbit the moon*

4 Analysis

Figure 3 depicts the Apollo spacecraft leaving Earth's parking orbit and crashing into the Moon, using two different changes in velocity. In order to do this the orbital mechanics of each body in

the system had to be taken into account. Once the arrival time was calculated it was possible to have the spacecraft orbit the Moon. This called for another instantaneous change in the velocity. Once a possible lunar orbital velocity is determined the task of brining Apollo 13 back to Earth is done with another instantaneous velocity change. If done correctly, this simulation will end with the safe return of all the onboard astronauts intact.

As seen in figure 4, it is no simple task to catch the orbit of the moon. Even after crashing into the moons surface it proved to difficult to have the spacecraft orbit the moon. Without the ability to wrap around the moon my program was not capable of returning the Apollo 13 rocket and its crew back to Earth.

5 Conclusion

Even after simplifications were made the instantaneous velocity corrections were not easy to determine. More often than not, the spacecraft would miss the Moon entirely. In some simulations the velocity correction was to small. This would have lead to the death of all the astronauts due to starvation and dehydration, as they drifted between Earth and the Moon. This simulation began with Apollo already in orbit. Having a rocket reach the Moon is no easy task. To get men from Earth to the Moon and safely back again is a task few have ever achieved.

References

- [1] R. H. Landau and M. J. Paez, "Computational Physics, Problem Solving with Computers," (Wiley: New York) 1997.
- [2] Gorham, Peter. "Physics 305 Differential Equations." P305lab7. Phys.hawaii.edu, 7 Apr. 2015. Web. 8 Apr. 2015.
- [3] "John F. Kennedy Space Center - Apollo 13." John F. Kennedy Space Center - Apollo 13. N.p., 4 June 2002. Web. 10 Apr. 2015.

Acknowledgements

All programs developed in modeling the motion of this projectile were written with the aid of Landau [1] and Gorham [2].